

Status of the McStas instrument simulation project

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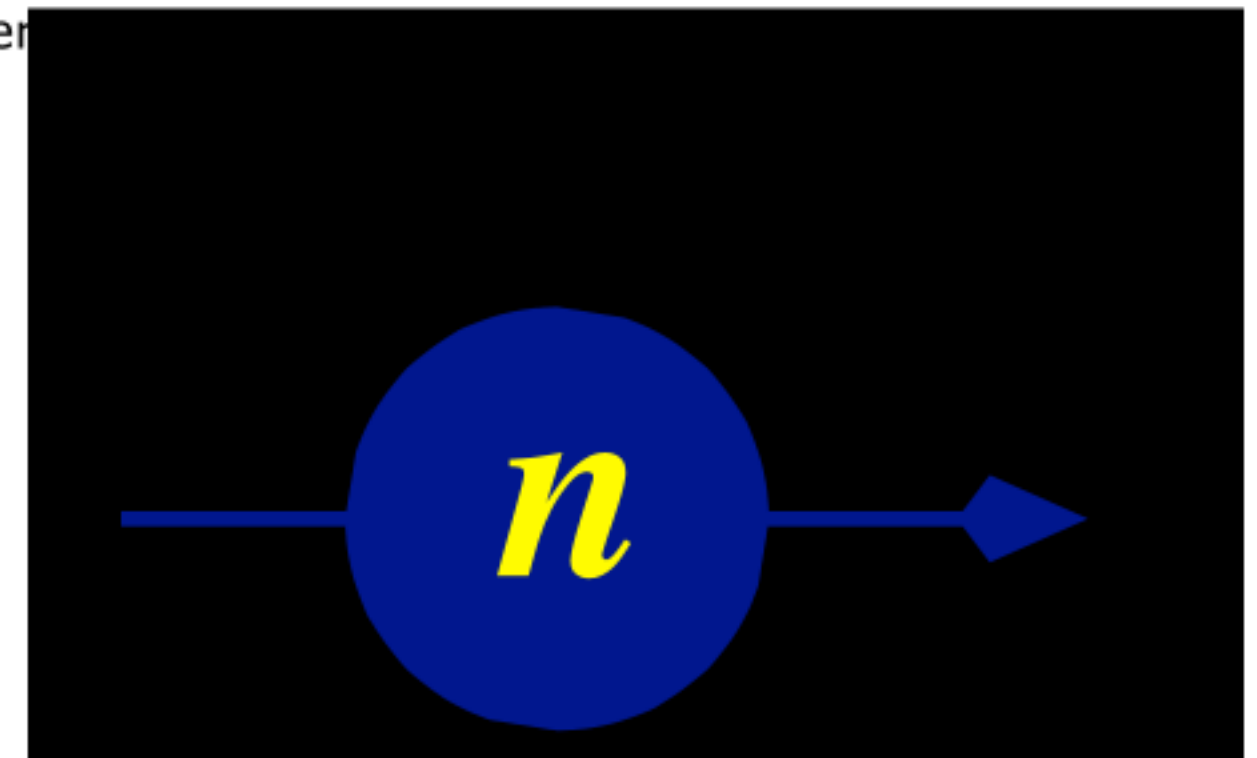
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⁴Niels Bohr Institute, University of Copenhagen, Denmark

⁵ESS design update programme, Denmark

⁶ESS design update programme, Switzerland



EUROPEAN
SPALLATION
SOURCE

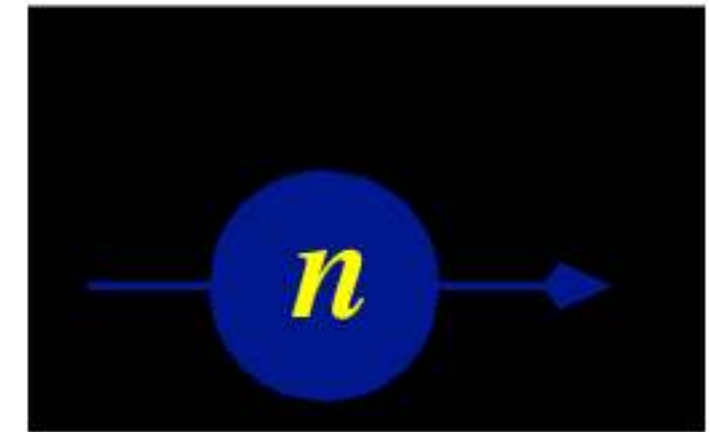
Status of the McStas instrument simulation project - NOBUGS 2012

Agenda

- An introduction to
 - McStas
 - Monte Carlo + raytracing
 - Project and technology
- New developments for the **long-awaited** 2.0 release
- Conclusions / lessons learned



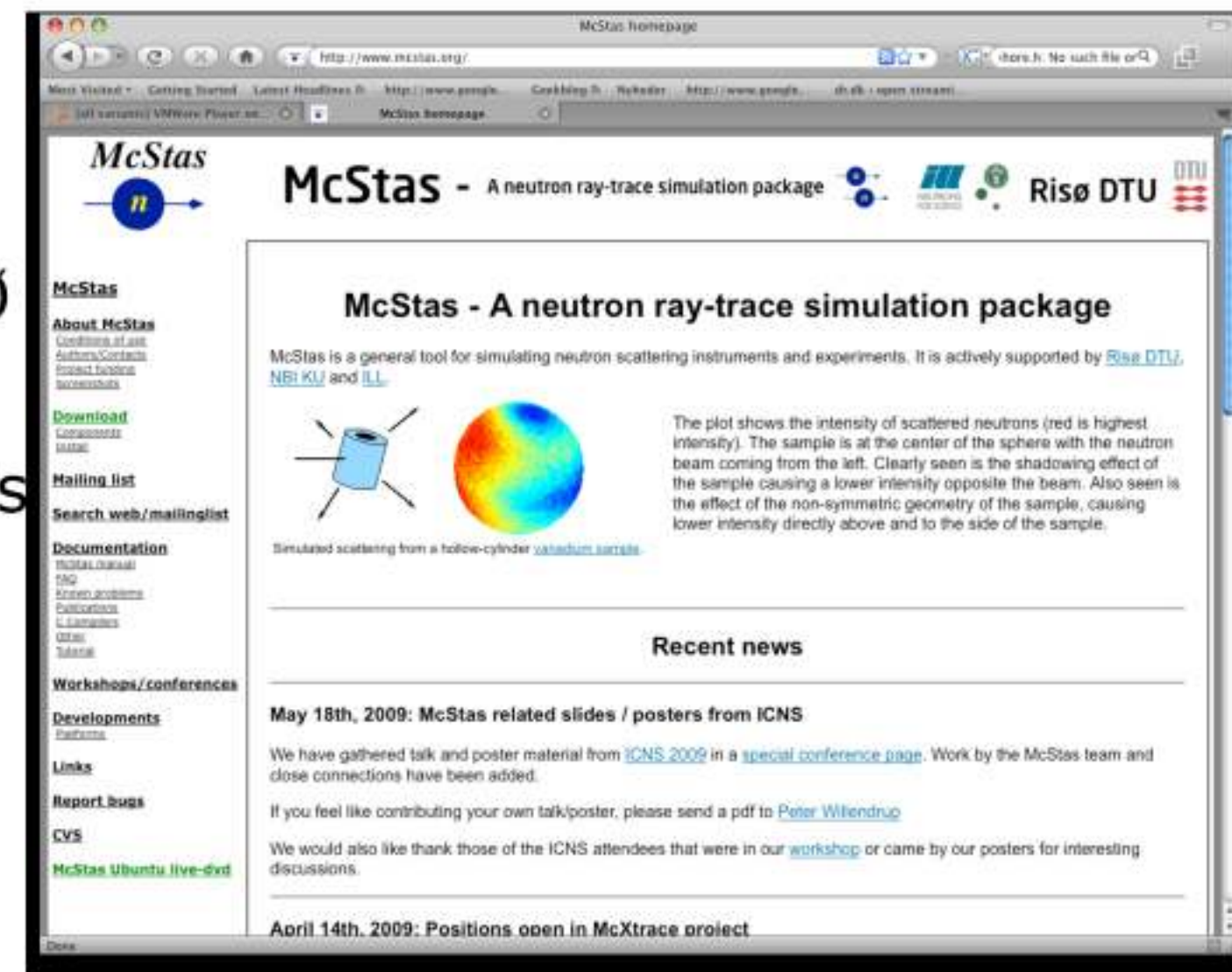
McStas Introduction



- Flexible, general simulation utility for neutron scattering experiments.
- Original design for Monte carlo Simulation of triple axis spectrometers
- Developed at DTU Physics, ILL, PSI, Uni CPH
- V. 1.0 by K Nielsen & K Lefmann (1998) RISØ
- Currently 2.5+1 people full time plus students



GNU GPL
license
Open Source



Project website at
<http://www.mcstas.org>

mcstas-users@mcstas.org mailinglist

See E. Knudsen poster!

McStas Introduction

McXtrace - since jan 2009 similar in X-rays

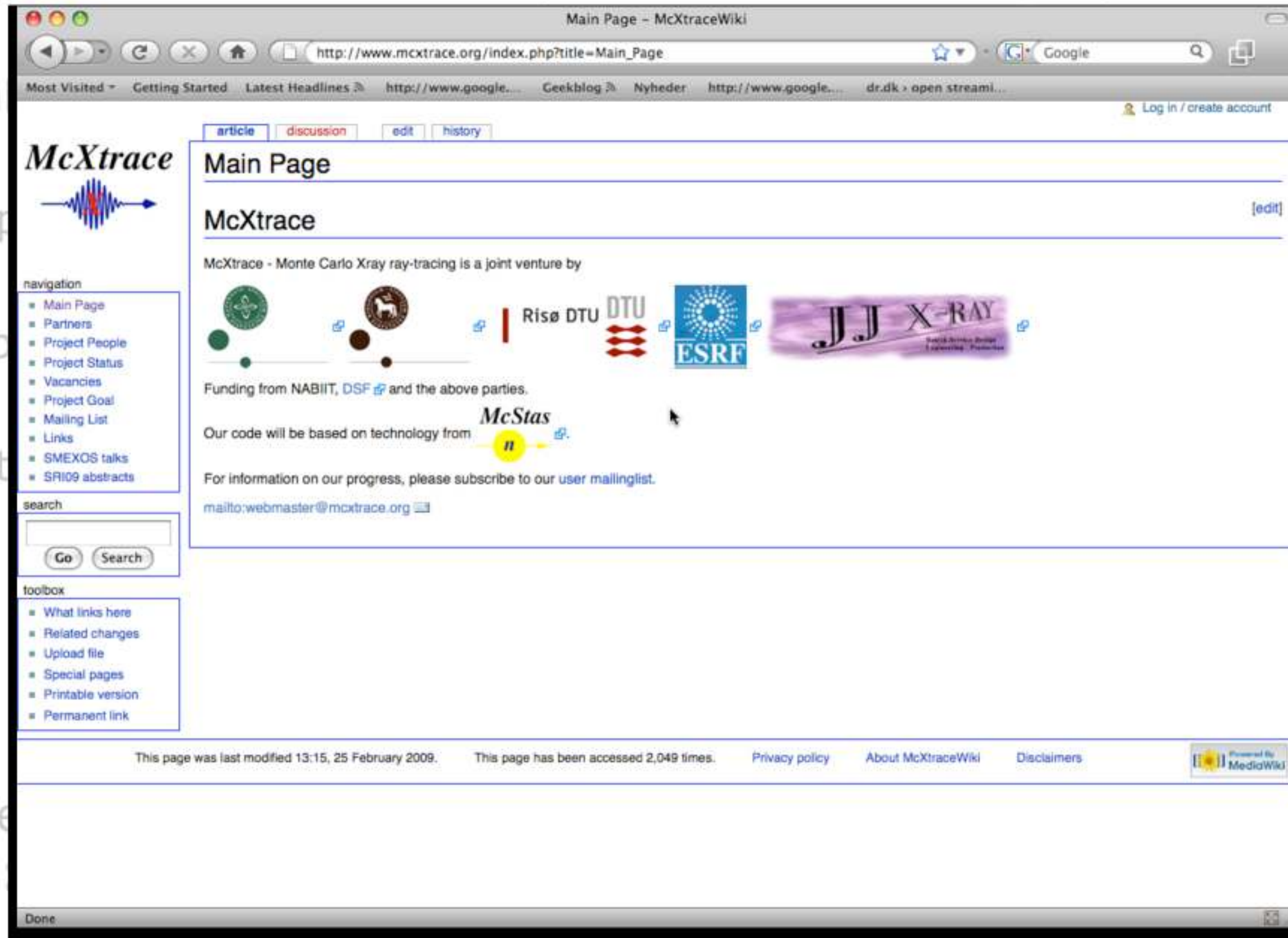
- Flexible, general simulation utility for neutron scattering experiments.

• Original

• Development

• V. 1.0 b

• Current

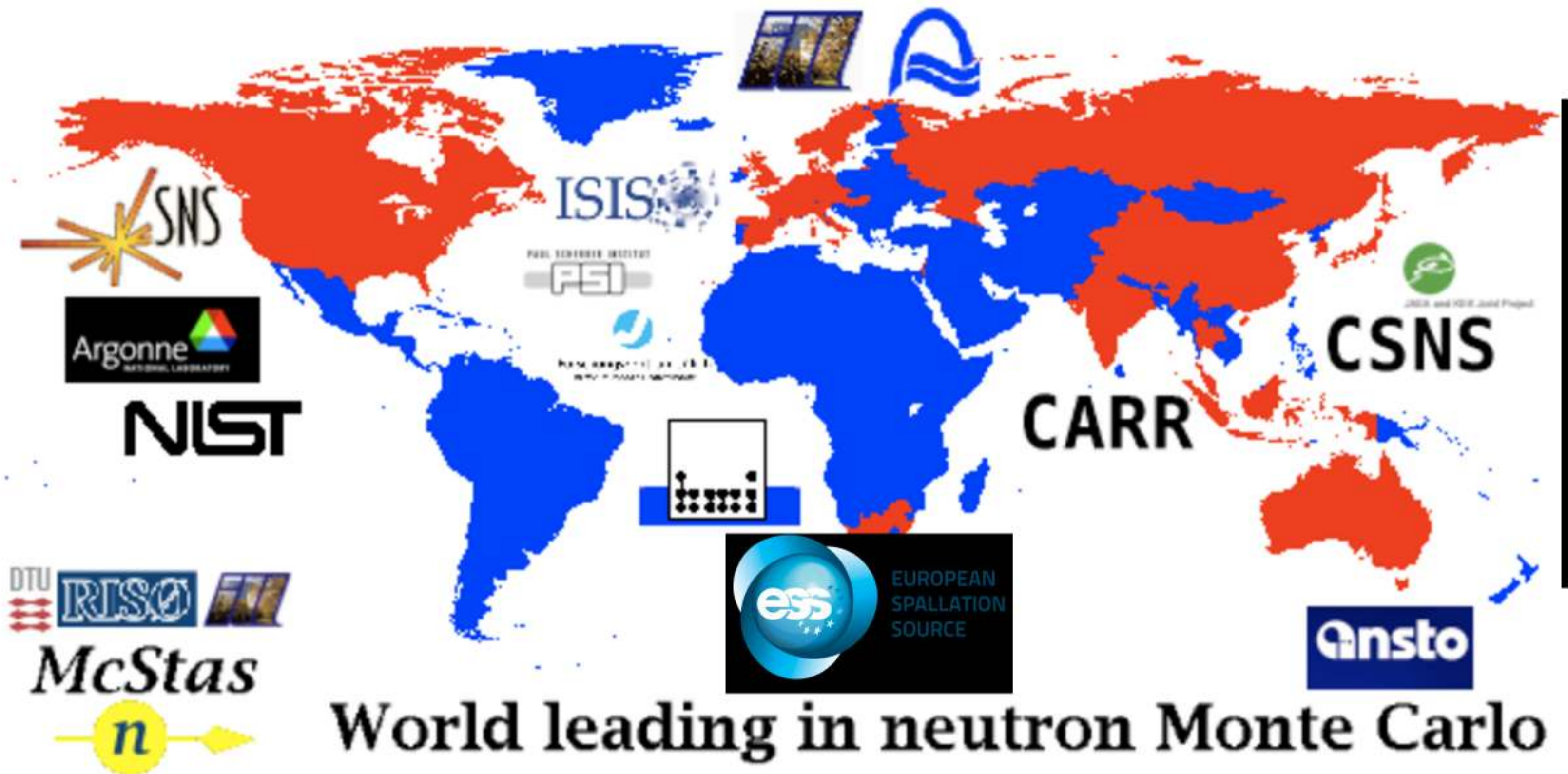


- Synergy, knowledge transfer, shared infrastructure

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McStas Introduction



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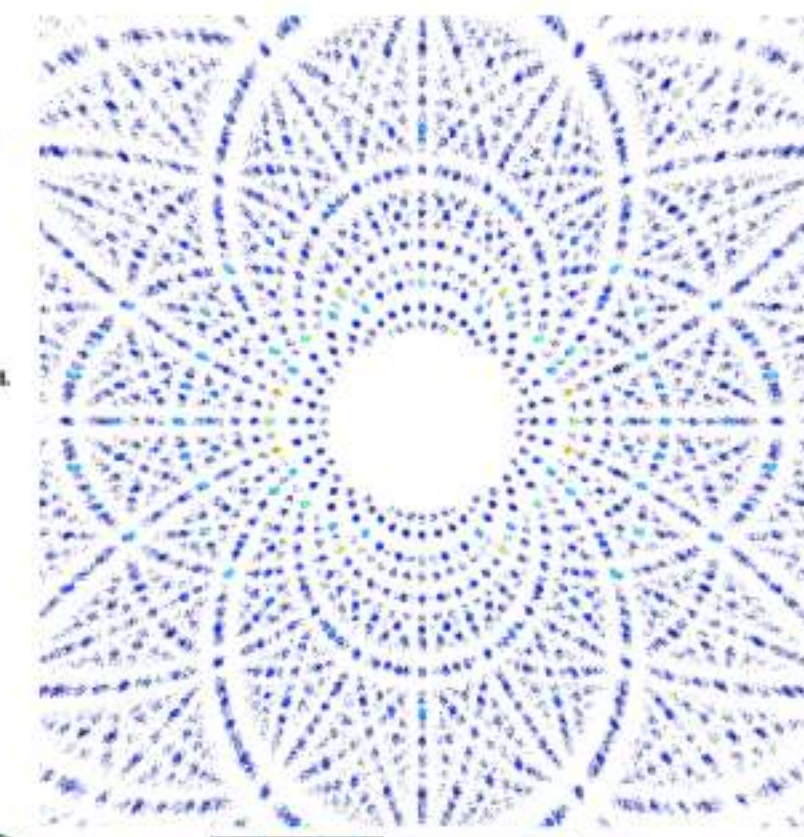
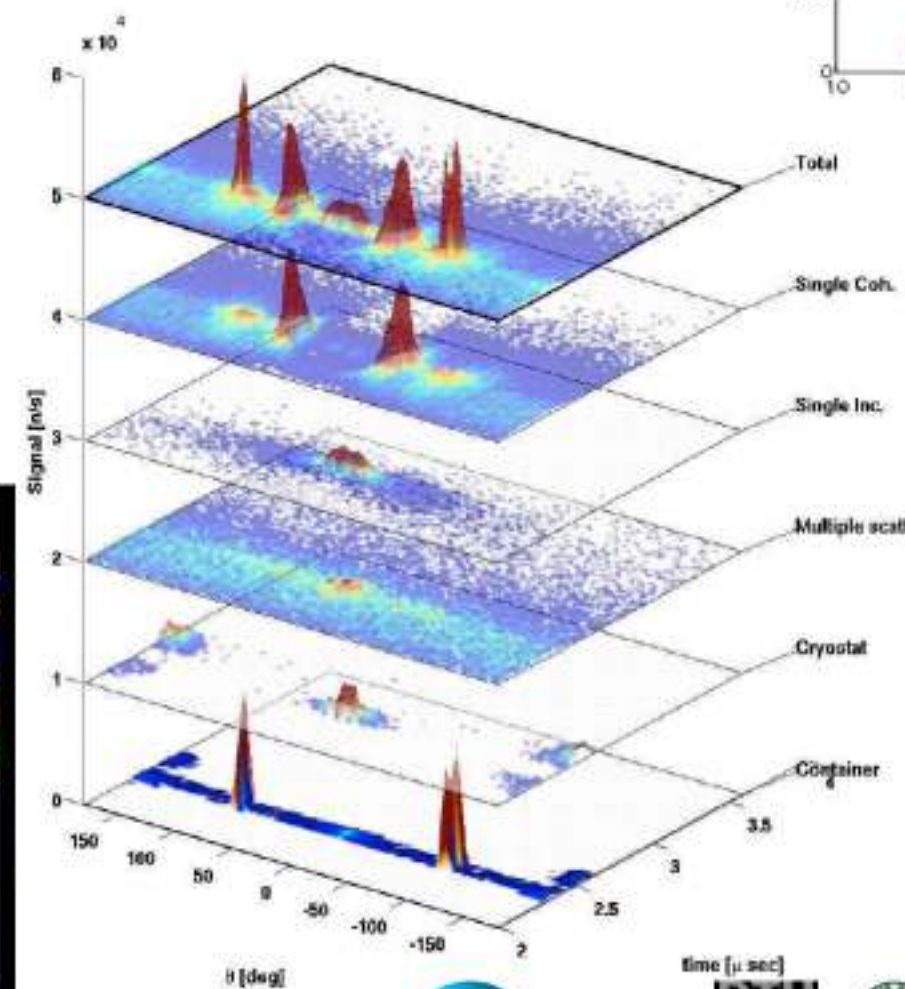
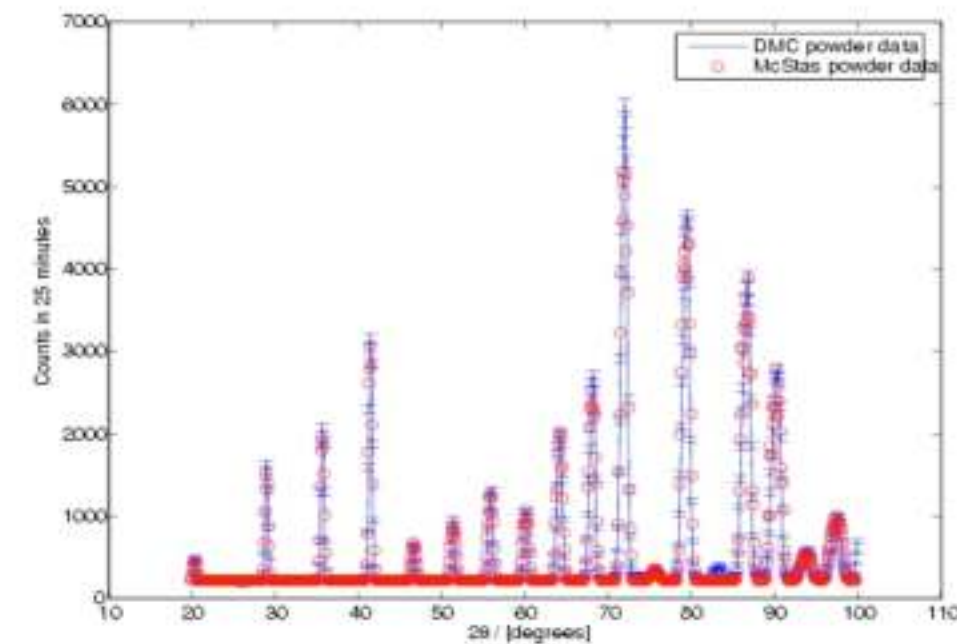
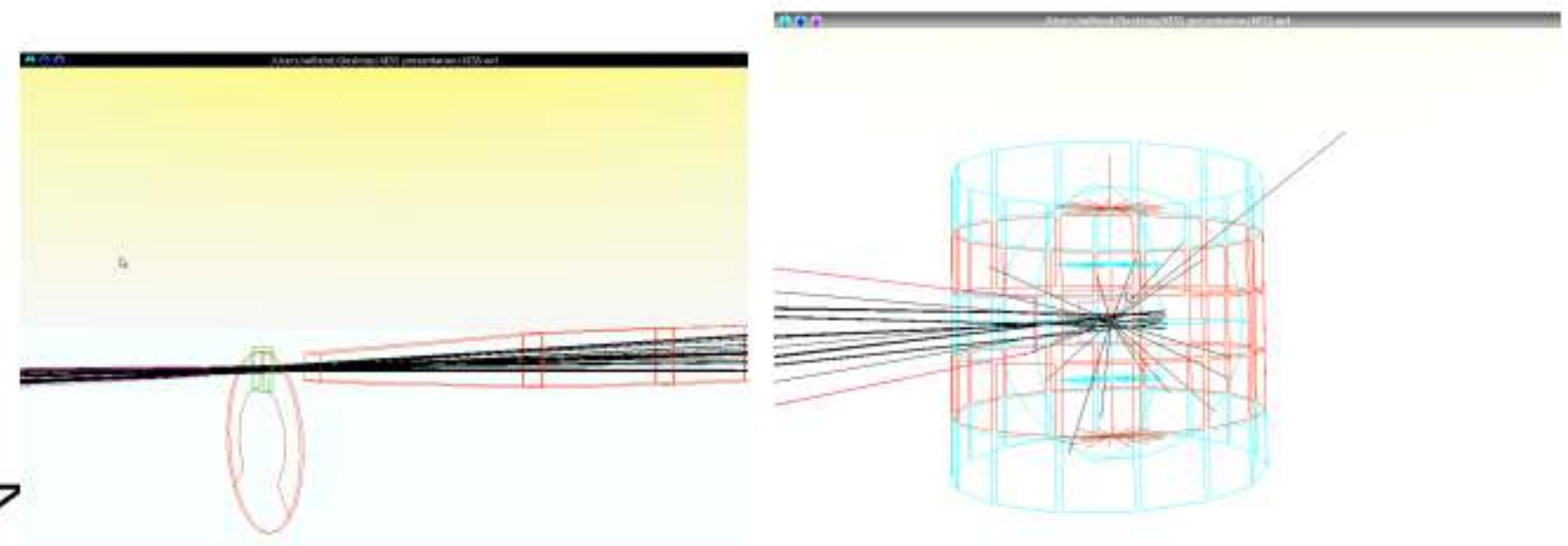
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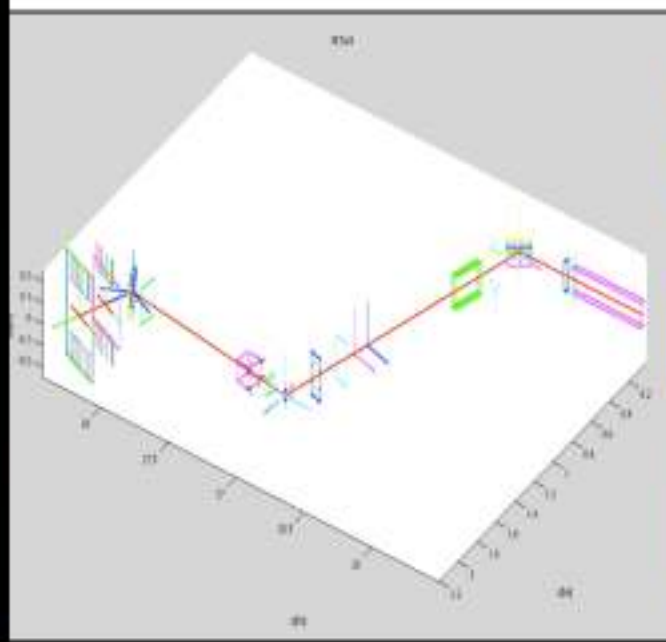
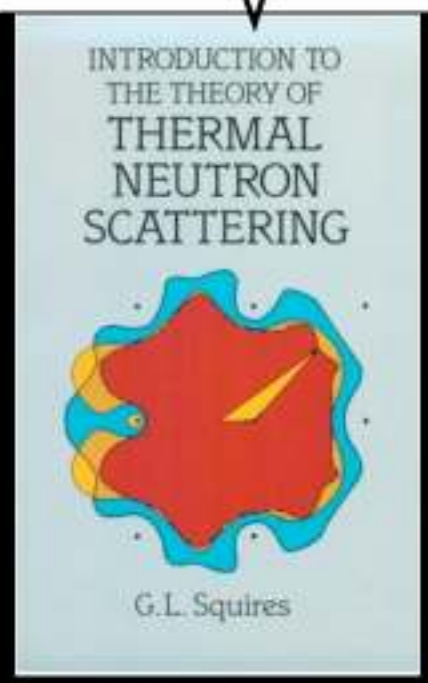
What is McStas used for?

- Instrumentation
- Virtual experiments
- Data analysis
- Teaching

KU, DTU 2005-2012
INSIS, NIDS, ESS workshops

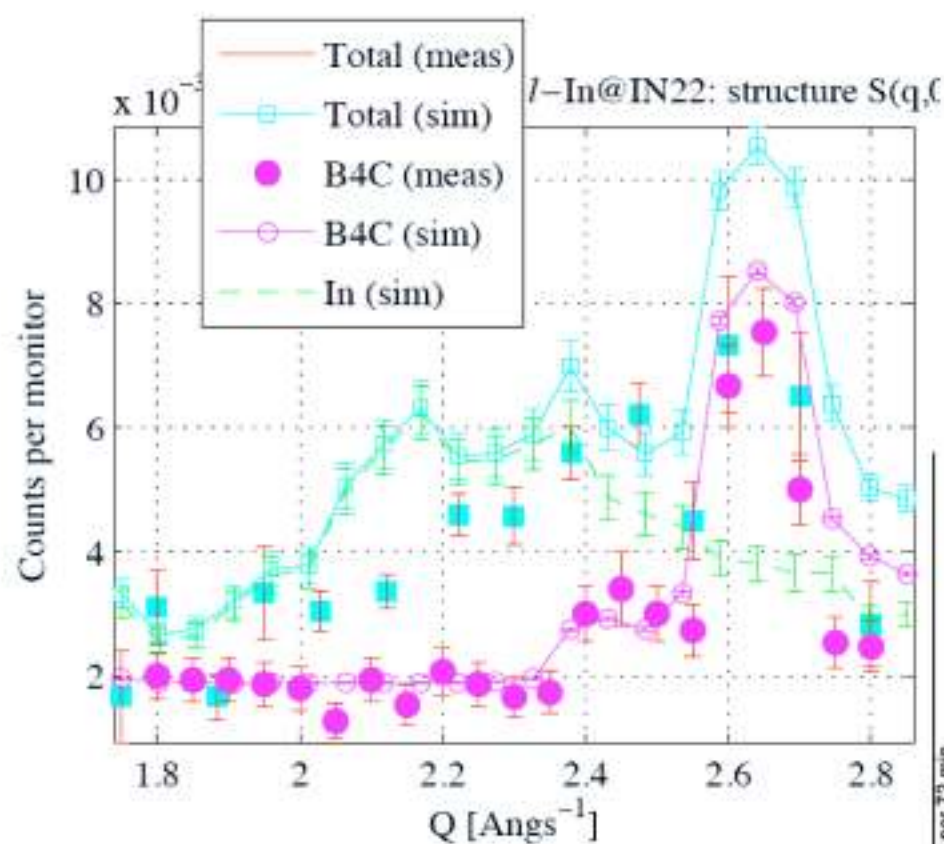


Simulation project - NOBUGS 2012

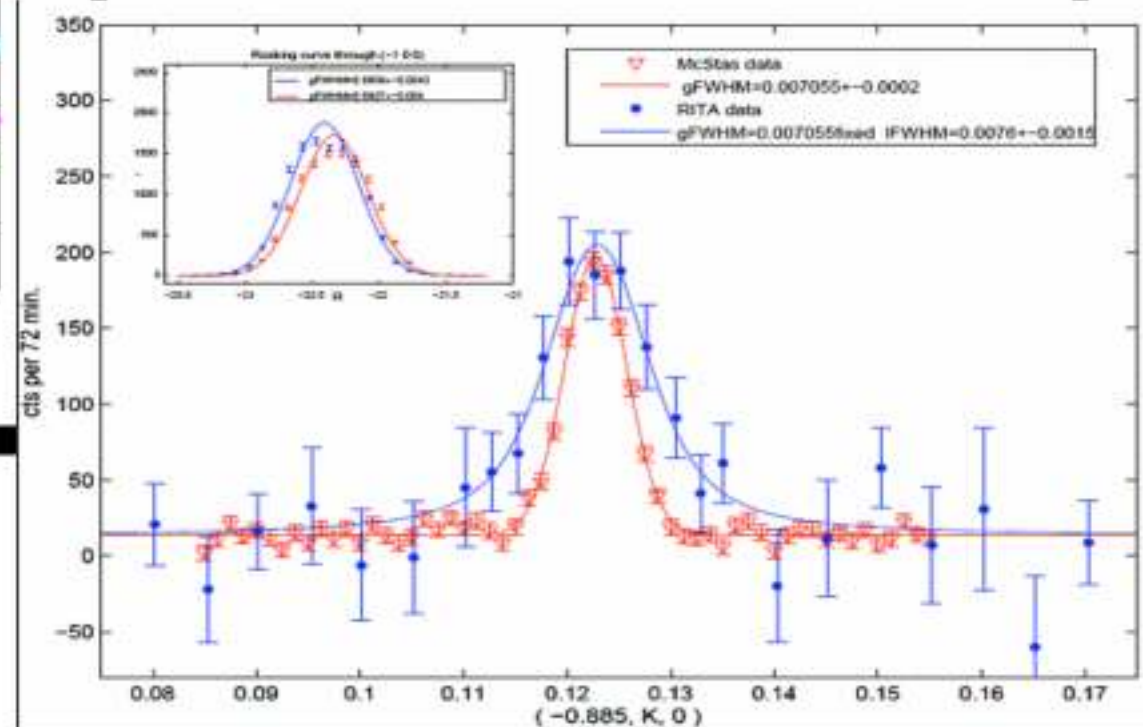


Reliability - cross comparisons

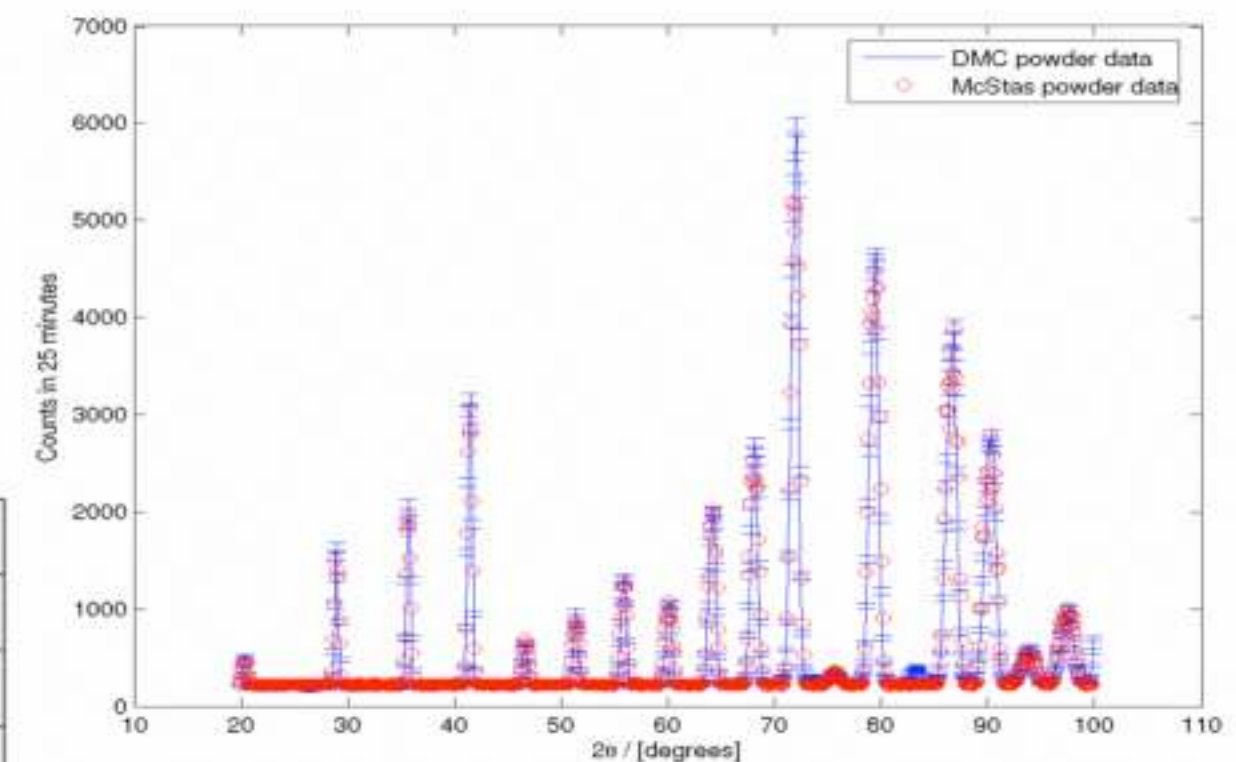
- Much effort has gone into this
- Here: simulations vs. exp. at powder diffract. DMC, PSI
- The bottom line is
- McStas agree very well with other packages (NISP, Vitess, IDEAS, RESTRAX, ...)
- Experimental line shapes are within 5%
- Absolute intensities are within 10%
- Common understanding: McStas and similar codes are reliable



E. Farhi, P. Willendrup et al., in



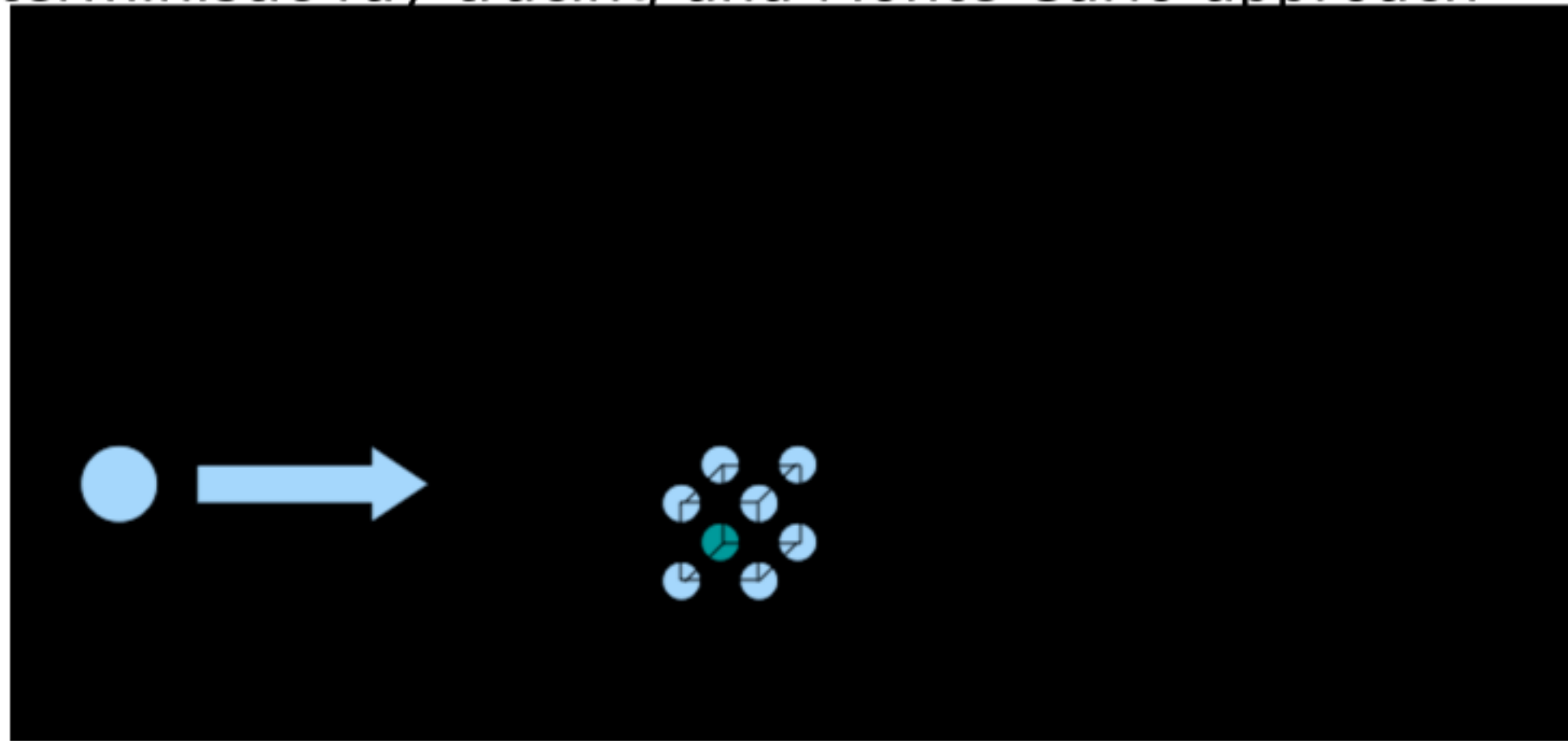
L. Udby, NIMA 2011, Volume 634 (2011)



P. Willendrup et al., Physica B, 386, (2006),

Elements of Monte-Carlo raytracing

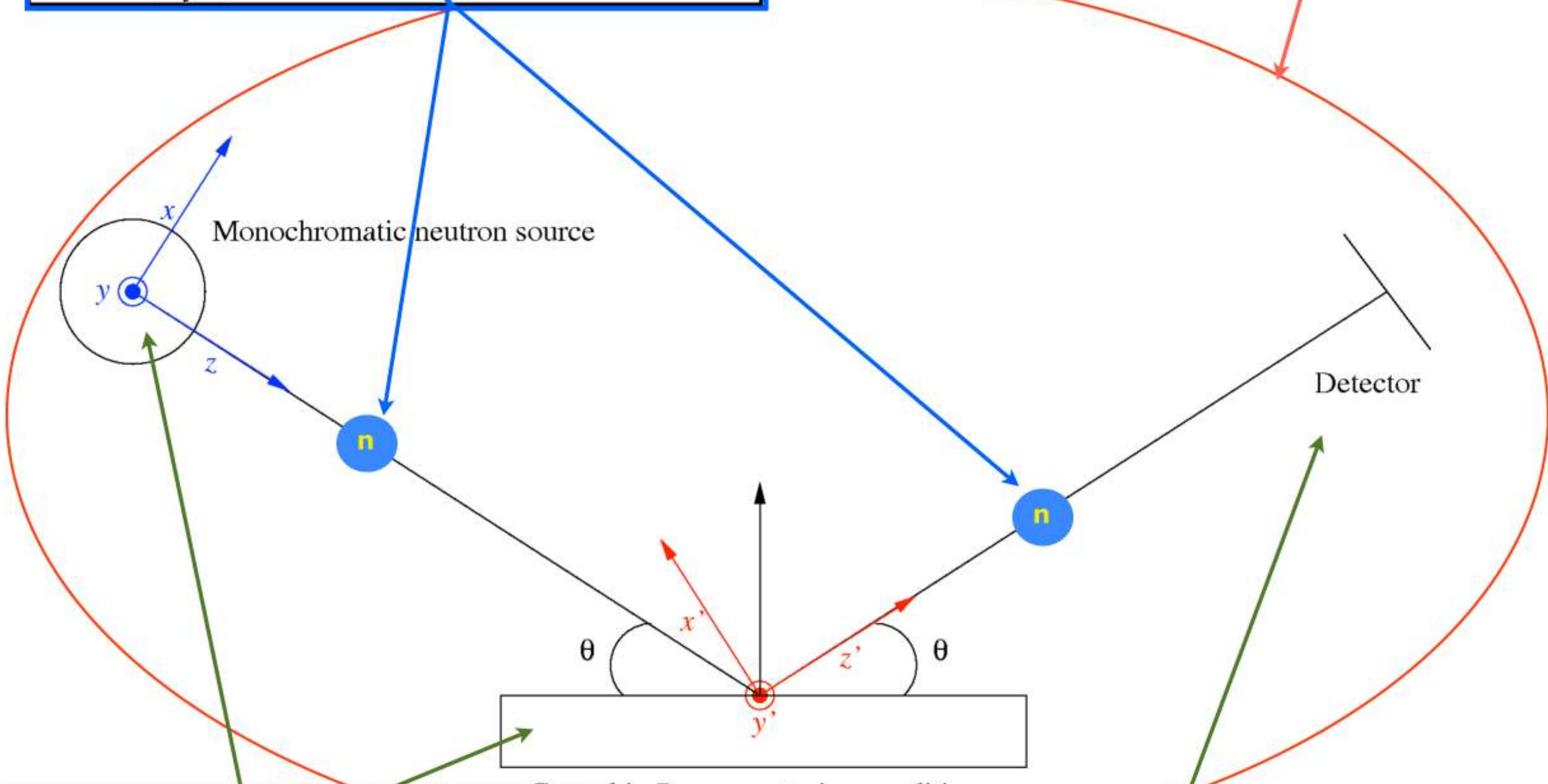
- Instrument Monte Carlo methods implement coherent scattering effects
- Uses deterministic propagation where this can be done
- Uses Monte Carlo sampling of “complicated” distributions and stochastic processes and multiple outcomes with known probabilities are involved
- - I.e. inside scattering matter
- Uses the particle-wave duality of the neutron to switch back and forward between deterministic ray tracing and Monte Carlo approach



- Result: A realistic and efficient transport of neutrons in the thermal and cold range

Neutron ray/package:
 Weight (p): # neutrons (left) in the package
 Coordinates (x,y,z)
 Velocity (v_x, v_y, v_z)
 Spin (s_x, s_y, s_z)
 Time (t)

Instrument: positioning + transformation between sequential component coordinate systems, e.g. neutron source, crystal, detector.



Components: Here the neutron physics happen, neutron weight adjusted according to scattering probabilities etc.

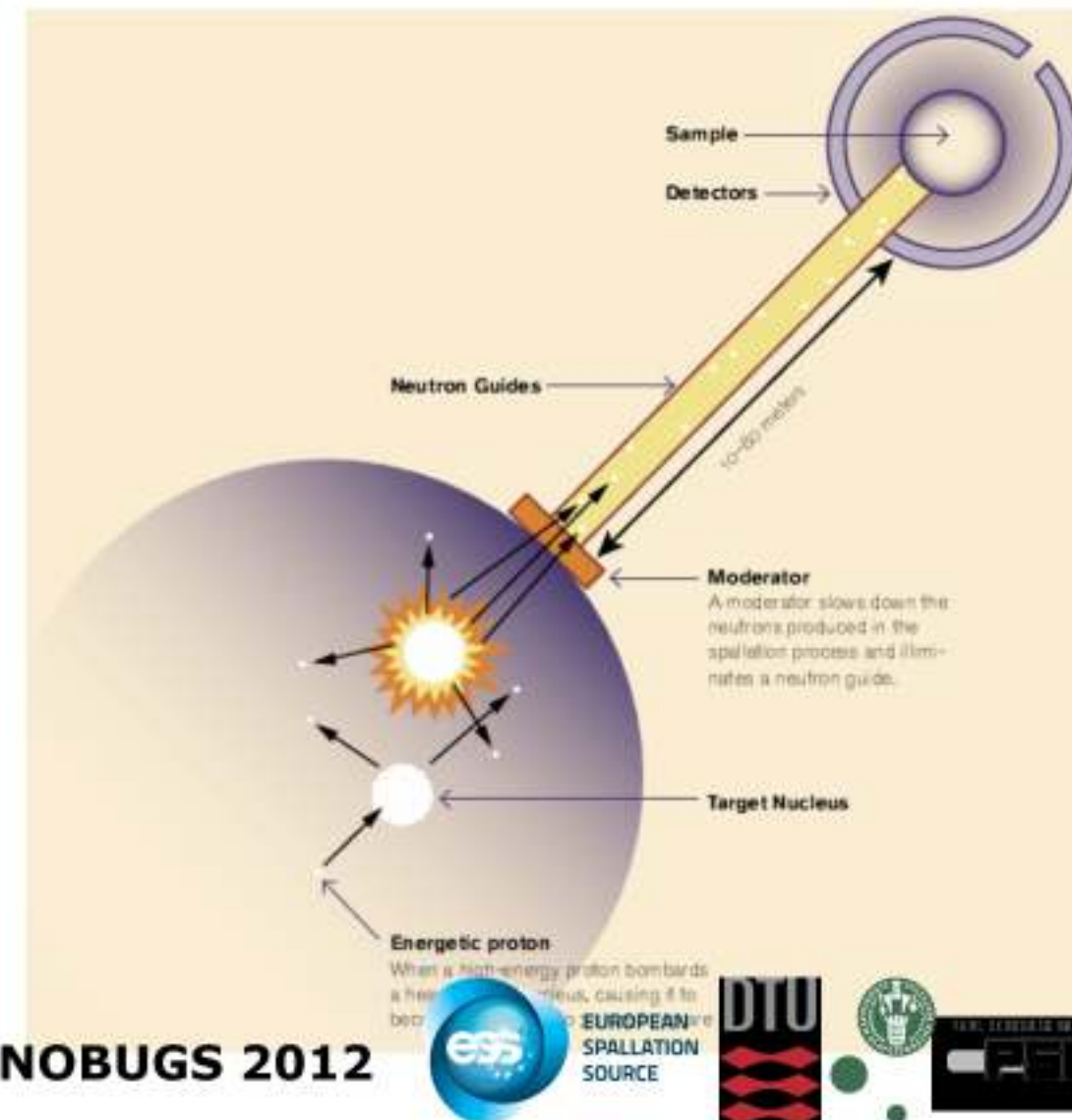
Local, internal coordinate system!

McStas overview

- Portable code (Unix/Linux/Mac/Windows)



- Ran on everything from iPhone to 1000+ node cluster!
- 'Component' files (~100) inserted from library
 - Sources
 - Optics
 - Samples
 - Monitors
 - If needed, write your own comps
- DSL + ISO-C code gen.



Under-the-hood / inner workings

- Domain-specific-language (DSL) based on compiler technology (LeX+Yacc)

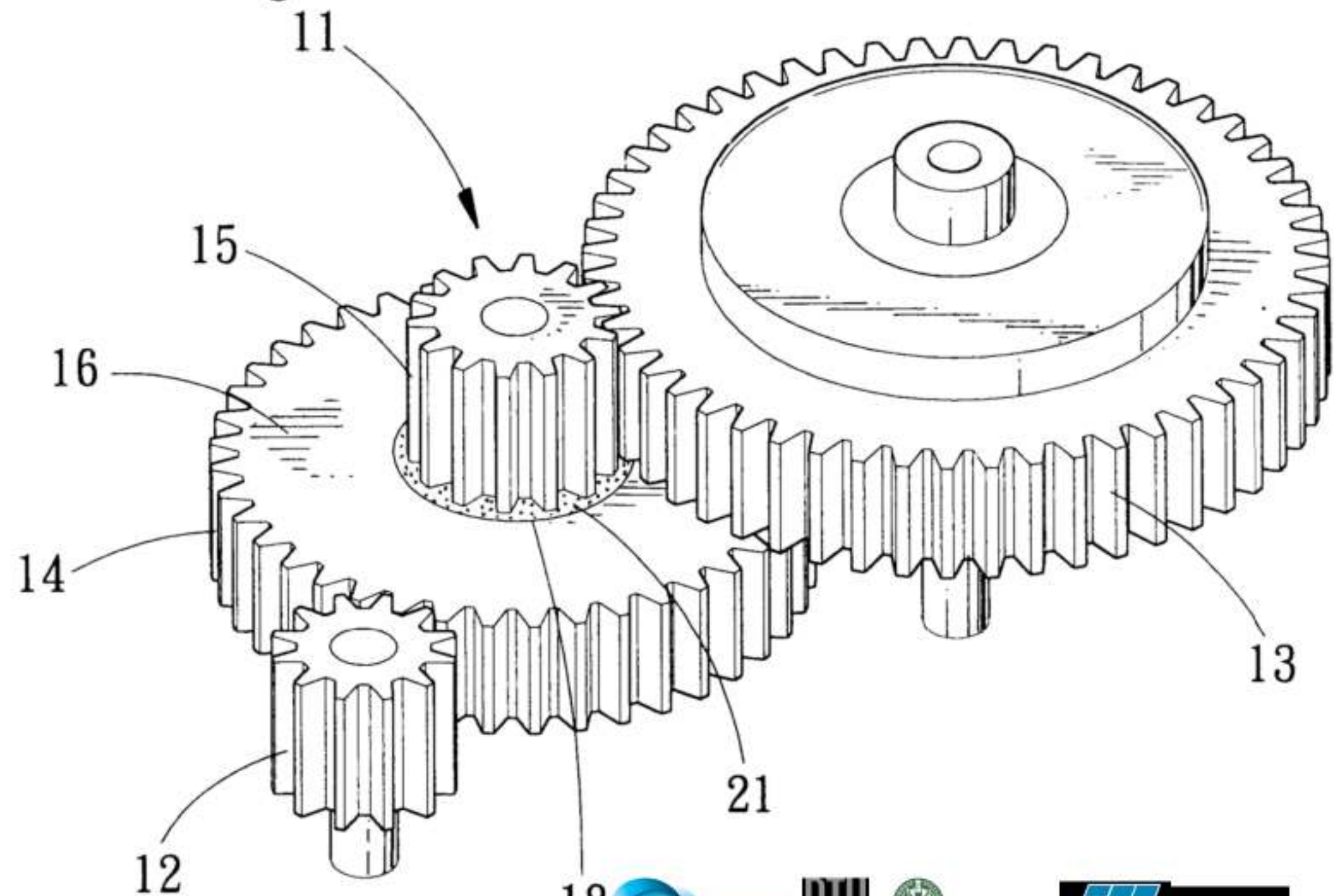
• Simple Instrument language $\xrightarrow{\text{Code generation}}$ ISO C



- Component codes realizing beamline parts (including user contribs)

- Library of common functions for e.g.

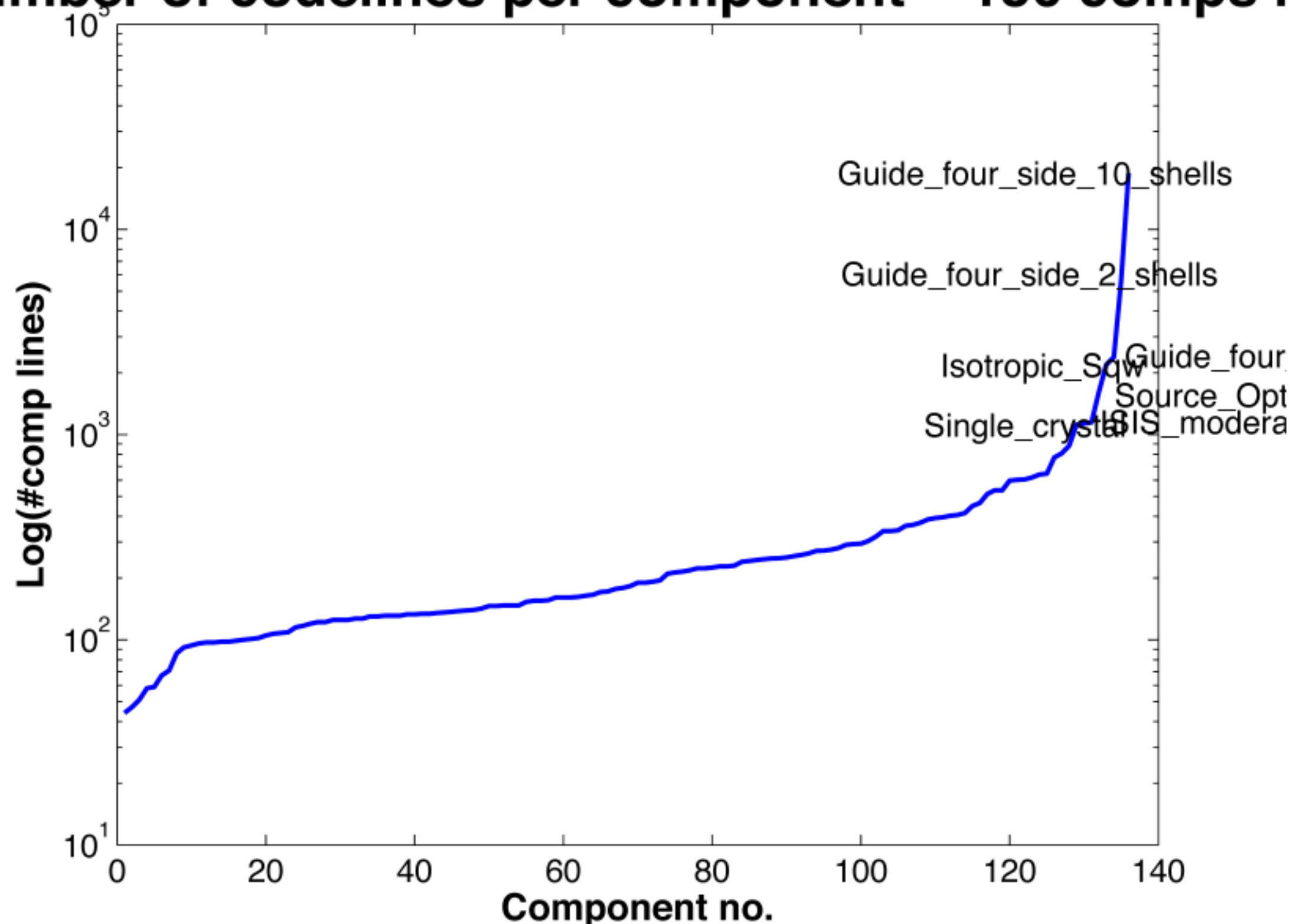
- I/O
- Random numbers
- Physical constants
- Propagation
- Precession in fields
- ...



Writing/modifying comps is not complex...

- Most comps are quite simple and short ~ 100 lines
- 30-40% of existing and new additions are from users
- Requirement: Test-scenario for documentation & unit test

Number of codelines per component – 136 comps i



Syntax in one, complex view...

```
{SPLIT} COMPONENT name = comp(parameters) {WHEN condition}  
  AT (...) [RELATIVE [reference|PREVIOUS] | ABSOLUTE]  
  {ROTATED {RELATIVE [reference|PREVIOUS] | ABSOLUTE} }  
  {GROUP group_name}  
  {EXTEND C_code}  
  {JUMP [reference|PREVIOUS|MYSELF|NEXT] [ITERATE number_of_times | WHEN  
condition] }
```



DECLARE / INITIALIZE

- Full flexibility of C in your instrumentfile!
- Use the DECLARE section define user variables and functions.
 - DECLARE %{\ul> - double myvar;
 - %}
- Use INITIALIZE for initialization of user variables and calculations.
 - INITIALIZE %{\ul> - myvar = sqrt(PI*input_var)*rand01();
 - %}
- - Both use normal c-syntax.
- BEWARE: (example) What you do in the c-style areas is c-standard, e.g. trigonometric functions from math.h use radians! - McStas placement specifiers work in degrees, etc...

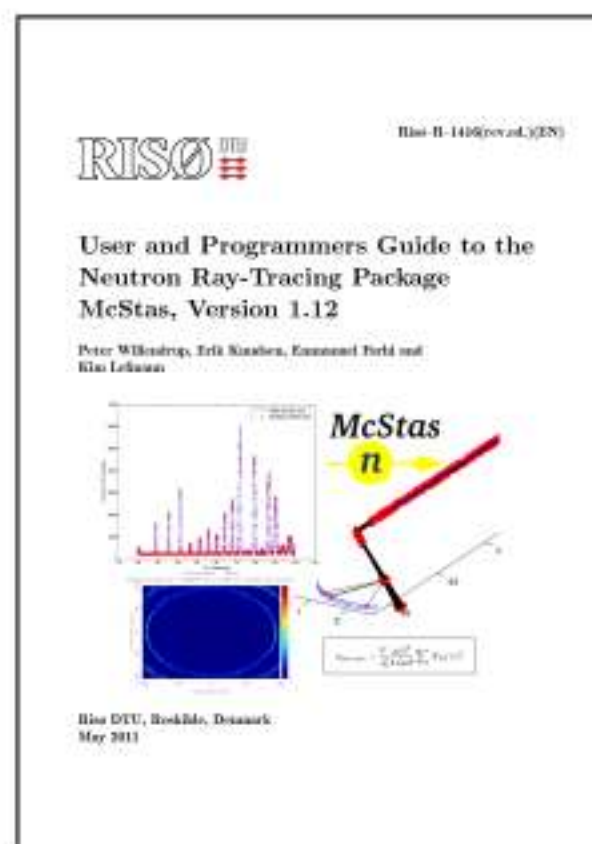


K & R



Documentation

- Basic use info is available inside comp & instr codes, extracted by perl to html
- 100+ page manuals documenting
 - Metalanguage
 - What is "under the hood"
 - Examples of practical use plus advanced features
 - Assumptions and algorithms applied in the components
- More than 70 example instruments
- Various tutorial and teach yourself solutions are available



McStas - Components/Instruments Library

File://A:/sets/willem/index.html

[sources | optics | samples | monitors | misc | control | examples]

[User Manual | Component Manual | McStas manual | Data files]

Components and Instruments from the Library for McStas

Names in **Boldface** denote components that are properly documented with comments in the source code.

Name	Origin	Author(s)	Source code	Description
Adapt_check	Risoe	Kristian Nielsen	comp	Optimization specifier for the Source_adapt component.
ESS_moderator_long	Risoe	KL, February 2001	comp	A parametrised pulsed source for modelling ESS long pulses.
ESS_moderator_short	Risoe	KL, February 2001	comp	A parametrised pulsed source for modelling ESS short pulses.
Moderator	Risoe	KN, M.Hagen	comp	A simple pulsed source for time-of-flight.
Monitor_Optimizer	J.L. (France)	Emmanuel Field	comp	To be used after the Source_Optimizer component
Source_Maxwell_3	Risoe	Kan Lefmann	comp	Source with up to three Maxwellian distributions
Source_Optimizer	J.L. (France)	Emmanuel Field	comp	A component that optimizes the neutron flux passing through the Source_Optimizer in order to have the maximum flux at the Monitor_Optimizer position.

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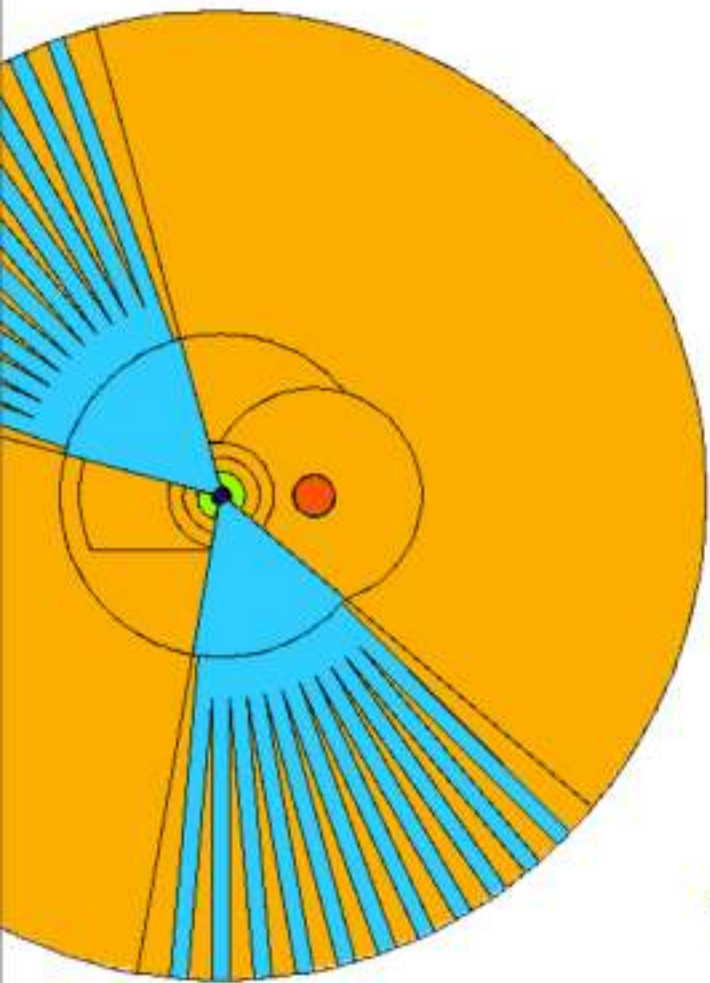


New developments 1 - work on McStas-MCNPX interfaces



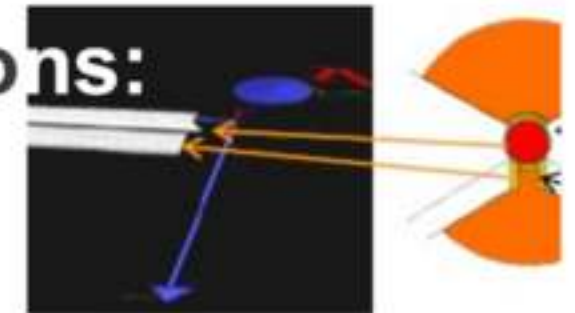
The task:

“Interfacing the MCNP and McStas Monte Carlo codes for improved optimization of the ESS moderator-beam extraction systems”



The solutions:

- Tally



- Ptrac

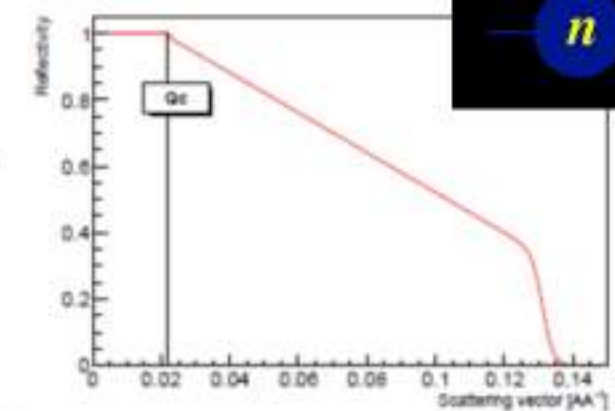
3000	2	10	179
100	2	0	
0.00000E+00	0.28640E+00		
0.43531E+00	-0.10000E+01		
0.00000E+00	0.00000E+00		
0.10000E+00	0.10000E+01		
0.33356E-02			

MCNPX
MCNPX

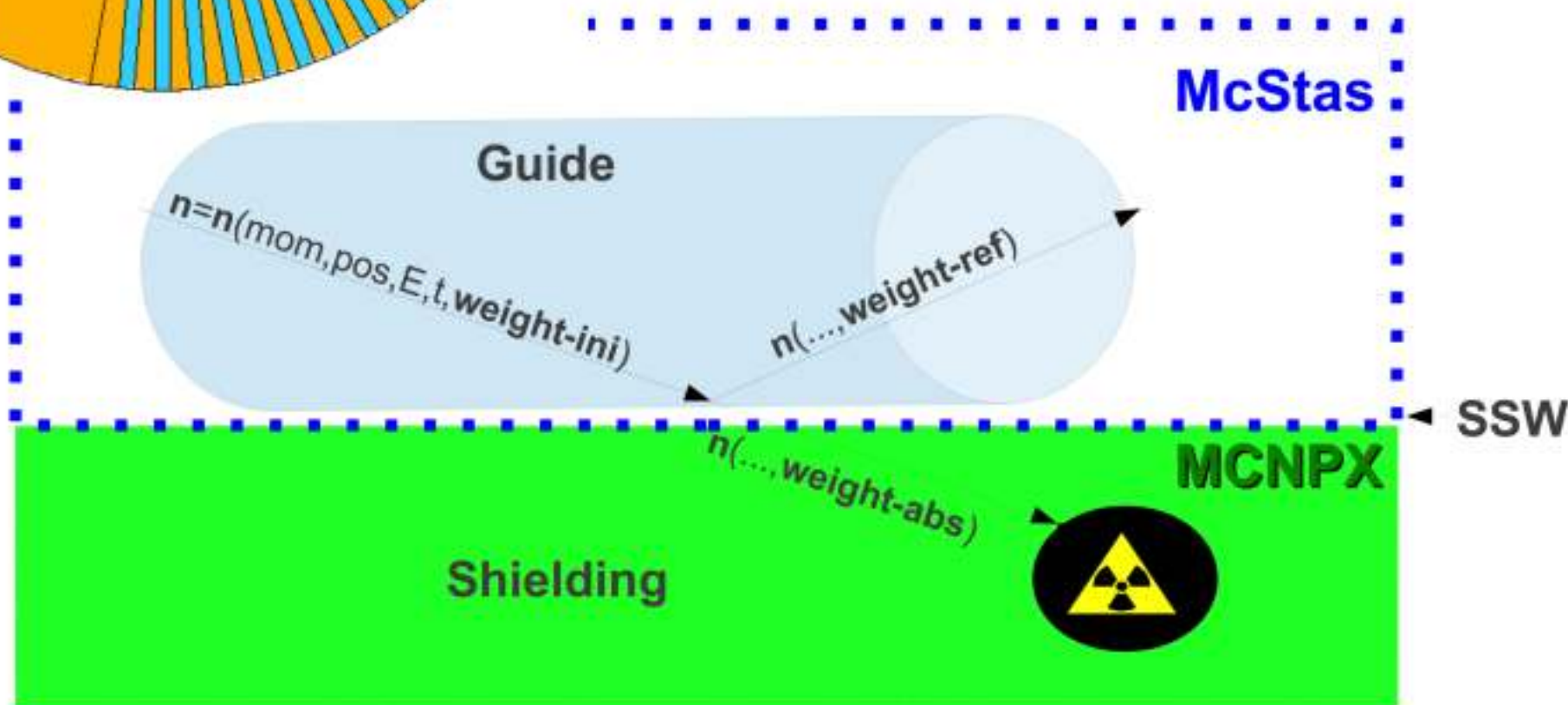
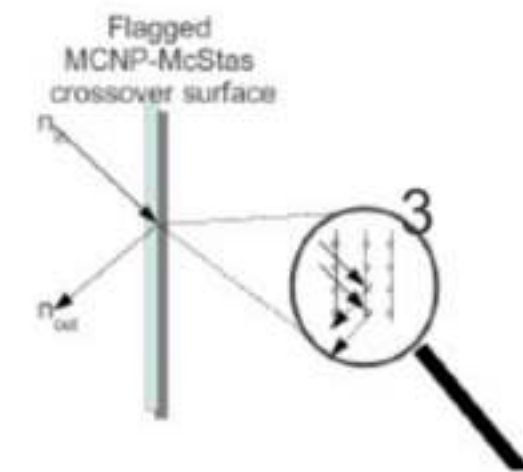
- SSW



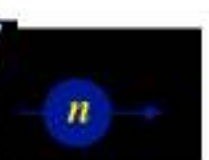
- Supermirror



- Compile

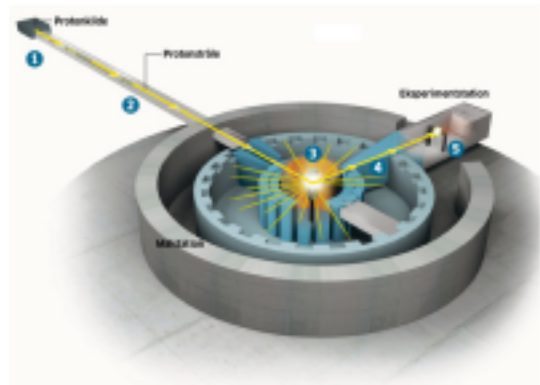


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New developments 2 - New Python-based tools and web frontend

See J. Brinch poster



Neutron simulations from the web

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1. Physics Department, Technical University of Denmark; {JSBN,PKWI}@FYSIK.DTU.DK

2. Institut Laue-Langevin, Grenoble, France

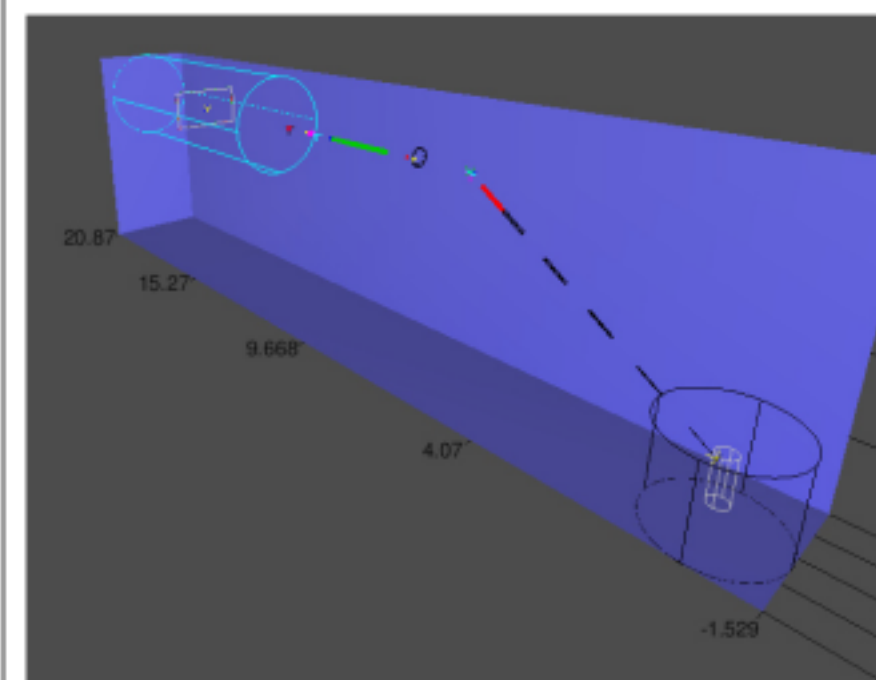
McStas



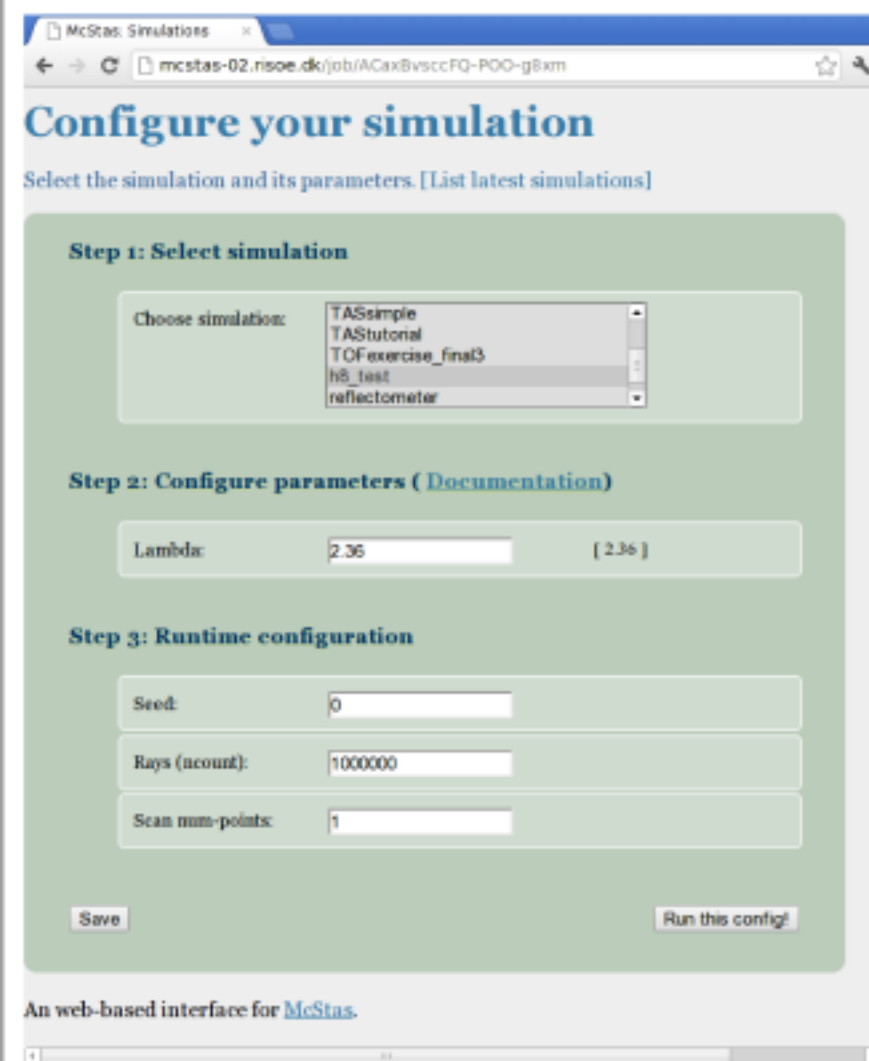
OVERVIEW

McStas is developing a web-interface for the simulation software, based on HTML and Javascript. The web-interface allows a user to run the McStas simulation software from a server; without installing it herself.

3D VISUALISATION



EXAMPLE: CONFIGURE



FEATURES

Permanent URLs: Both the URLs for configuration and results are unique and permanent; they can easily be shared with fellow students or colleagues.

Parameter verification: A sanity check eliminates basic mistakes before running the simulation. Most helpful for students.

Authentication: A simple authentication measure prevents anonymous users from using the system.

Limits: Upper limits can be set for the neutron count to prevent users from running very long simulations.

Parallelism: Simulations can be run in parallel; both with many workers and on a per simulation basis using MPI.

1. CONFIGURE

First, the instrument to simulate is chosen and configured. Parameters are adjusted and saved before the simulation is started.

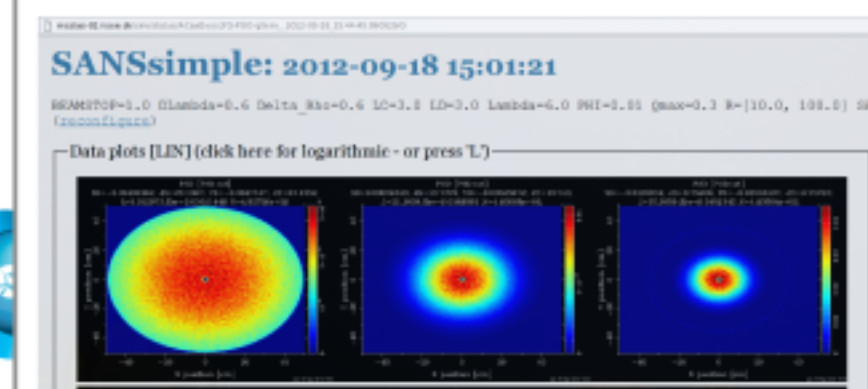
2. SIMULATE

Once the simulation has been configured, it

EXAMPLE: RESULT OVERVIEW

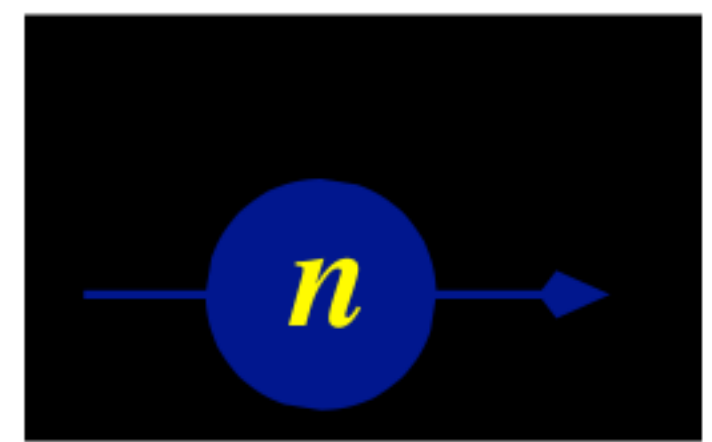


SCANNING A PARAMETER



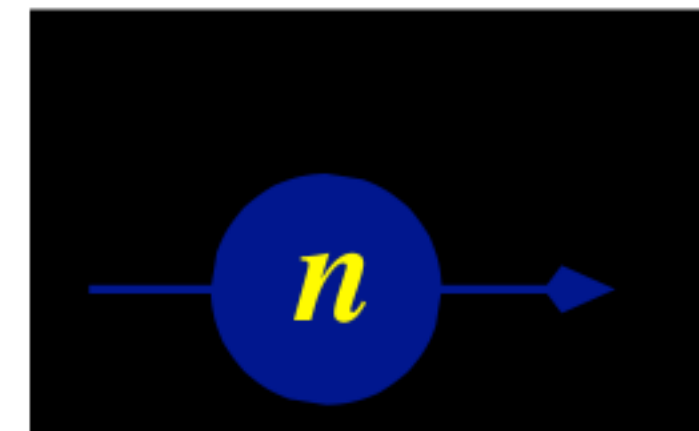
Other McStas 2.0 highlights

- Standardized parameter naming
- Updated library of the sources - common effort with Viteess
 - recommendation documents
- Split GUI/tool and calculation layer installations
 - mcstas-core, mcstas-comps, mcstas-tools-perl, mcstas-tools-python, ...
- Proper 64 support incl. MPI on Win64 and OS X
- New contributions in various areas
- Improved support for polarized neutrons
 - "All" components support polarized neutrons
 - Precession in tabulated magnetic fields
- Release in October 2012 (!!) - feel free to try the current dev tree
 - ~# svn co <https://svn.mccode.org/svn/McCode>
 - ~# ./configure --enable-mcstas



Conclusions incl. lessons learned

- McStas is a stable project and code
- Open Source strategy and modular design part of reasons for success
- Start out with healthy design
 - Original design by computer scientist, allowed physicists to expand like hell! Still rock solid! :-)
- Choose simple over complex or at least provide both options
- Flexibility is important - users are different
- Documentation is of key importance
- First user experience must be that "it works"
- Beyond 2.0 we will be applying "release often, release early..."



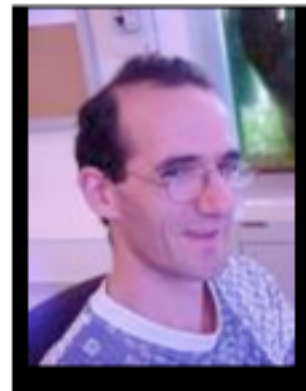
People

- The success of the project is also about the people:

- Present McStas team members



• K Lefmann



E Farhi



P Willendrup



E Knudsen



U Filges



J Brinch

- Past McStas team members



• K Nielsen



PO Åstrand



K Lieutenant



P Christiansen

